

Physics of Galaxies 2017 Lecture 1: Introduction



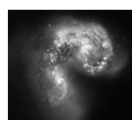
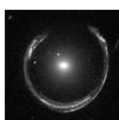
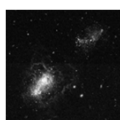
Outline for today I

- Formal Stuff:
 - Course literature
 - Examination
 - Schedule
- Course outline



Outline for today II

- What is a Galaxy?
- Historical Background
- Galaxy Classification
- The Cosmological Framework



Teacher

- Erik Zackrisson
 - Email: erik.zackrisson@physics.uu.se
 - Room 63103
In astronomy corridor on floor 3 in house 6
– just ring the bell to get in!

Course homepage

- Link:
www.astro.uu.se/~ez/kurs/Galaxies17.html

The Physics of Galaxies, 10 ECTS (3p), Spring 2016
 Literature: Peter Schneider, 2015, "Extragalactic Astronomy and Cosmology", Springer, ISBN 978-3-642-54082-0 (hardback) or 978-3-642-54083-7 (ebook)
 Contents: Extragalactic astronomy, with emphasis on the properties, evolution and origin of galaxies.
 Prerequisites: Knowledge corresponding to a Bachelor's degree in physics, or similar. In addition, basic knowledge about space, stellar physics, galaxies and cosmology corresponding to the course "Astronomy 1" is required.
 Examination: Seminars, literature report, hand-in exercises, laboratory exercise
 Teacher: Erik Zackrisson, erik.zackrisson@physics.uu.se
 Time: April-June 2016
 If you are interested in participating, please notify Erik Zackrisson as soon as possible!

Schedule

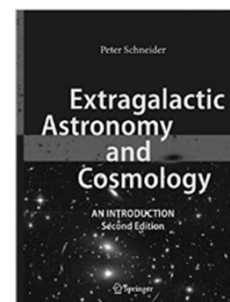
Recent changes are marked in red

Date	Time	Room	Lecture/ Exercise/seminar	Topics	To read	To turn in/prepare
April 12	16-17	A40115	Lecture 1	Course introduction Historical background The cosmological distance scale Galaxy classification	0-1.4, 1-3.1.2, 3.8-3.9.6	
April 18	15-16	A2003	Lecture 2	The Milky Way The Local Group	2.1-2.4.2, 4.1-4.1.3	
April 19	15-17	A40109	Lecture 3	Dark matter in galaxies	2.4.3, 3.3.3, 4.4.6, 7.6-7.6.3, 7.6, 8.2.2-8.2.3	

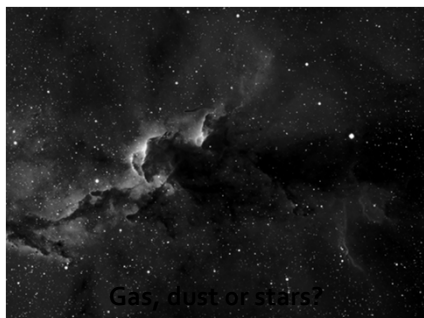
Course literature

Extragalactic Astronomy and Cosmology
 Peter Schneider
 2014/2015, Springer
 Hardback:
 ISBN 978-3-642-54082-0
 eBook:
 ISBN 978-3-642-54083-7

Around 700 SEK
**Note: E-version available
 from UU library**



Intermission: What are you looking at?



Examination

- Three exercise sessions
- Hand-in exercises
- Three seminars
- One computer/laboratory exercise
- Written essay (minimum 3 pages) + oral presentation (10 minutes)

But no written test!

Exercise sessions

- **Session 1:** April 20, 13-15
- **Session 2:** April 27, 13-15
- **Session 3:** May 17, 10-12
- **Objective:** Solve problems *together* in class



Exercise sessions

- **Preparation:**
 - Bring pen, paper, calculator/computer, preferably textbook
 - Session I: No preparation required
 - Session II/III: Study exercises and solutions posted on course homepage
- **Grade:** Pass/Fail
 - No-show or not actively participating →
 - Need to complete more hand-in exercises

Exercise session I: Fermi problems

Objective: Gain skill in making back-of-the-envelope calculations

Examples:

- How much gold is there in the Milky Way?
- How many galaxies are there in the Observable Universe?



Compete in teams - win marvelous prizes!

Exercises and solutions on the course homepage

Make sure you understand the solutions before coming to exercise session II & III!

The problems we solve in class will be similar.

2. **Brightness and colors of unresolved stellar systems.** Two stars in a close binary system have magnitudes 14.2 and 14.5, respectively. The first star has surface $T_1 = 4200$ K and the second $T_2 = 3500$ K. If this system is observed in a wavelength which cannot resolve the two components, what would the integrated flux and color $J - K$ of this object be?

Solution: At large distances, it becomes increasingly difficult to study the individual stars of stellar populations. They are usually too numerous for this.

1. The flux sensitivity of your telescope may be insufficient to allow the detection of a single, very distant star, whereas the combined light from large numbers of stars at the same distance may push them above the detection limit.

2. Limitations in angular resolution of your telescope may blend the light from nearby stars, thereby making distant stellar populations appear as single objects.

Black of contemporary extragalactic astronomy is founded almost to the study of the integrated light from large numbers of stars. To illustrate this, we show the appearance of stellar populations at increasing distances, where individual stars cannot be resolved.

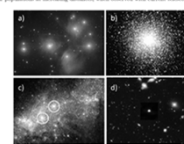


Figure 1: Resolved versus unresolved stellar populations. (a) A nearby star cluster (the Pleiades, distance 440 ly) in which individual stars can be resolved, by a Milky Way globular cluster (27 Mpc away).

Hand-in exercises

- 3 exercises downloadable from the course homepage
- Submit by email
- **Deadline:** June 14
- **Grade:** Fail, 3, 4, 5
- Collaboration OK, but please don't turn in identical solutions!

Note: If you didn't actively participate in the exercise sessions, you need to hand in additional exercises – please contact me if this situation should arise

Physics of Galaxies
Hand-in exercises 2017

Instructions: There are three problems you need to solve to pass the course. If you have already actively participated in the three exercise sessions, do not need to participate in order of them, you should submit the number for additional problems to solve in order to pass the course. Most exercise solutions are quite complex, but solutions received in fully completed and will also be awarded a grade, as these exercise sessions are intended to help you learn about the topics and encourage you to continue learning. The deadline for handing in solutions to these problems is June 14, 2017.

1. Cluster star formation. How many stars are there in the observable Universe? Make an order of magnitude estimate of this, quantify the uncertainty and make a table of the most important astrophysical quantities that are likely to affect your estimate (and clearly explain why they do so).

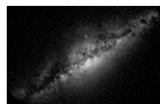
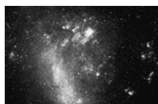
2. Population synthesis. Use the table of stellar parameters below to generate a simple population synthesis model. Assume that the stellar population of your model galaxy only consists of these types of stars (G, K, M, all kinds of the same size, and that the relative number of stars of each type is given by the numbers 100, 10, and 1).

Intermission: What are you looking at?



Literature exercise

- Choose subject individually
- Find suitable articles
 - Published papers (ADS abstract service) http://adsabs.harvard.edu/abstract_service.html
 - Preprints: <http://www.arxiv.org>
- Written report (≥ 3 pages), deadline May 24
 - **Grade:** Fail, 3, 4, 5
- Oral presentation (≈ 10 minutes), May 30
 - **Grade:** Fail, 3, 4, 5

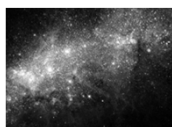
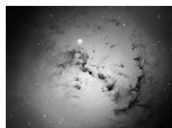


Required format of written report

- Abstract
- Introduction
- Main text (with references)
- Reference list
 - Should be mostly research or review papers
 - Please avoid using the textbook, popular science papers or homepages as references
 - Exception:** Links to project pages of upcoming telescopes, surveys etc. may be necessary if there is no proper paper out yet

Suggested topics

- The first stars
- Origin of supermassive black holes
- Galaxies and cosmic reionization
- Ultrafaint dwarfs
- Extragalactic background radiation
- Galactic archeology
- Magnetic fields in galaxies
- Conditions for life on galactic scales
- Science cases of future telescopes (pick one!):
 - James Webb Space Telescope
 - European Extremely Large Telescope
 - Square Kilometer Array



But please feel free to suggest other topics!

Seminars

- Small "simulations" of what the working as a scientist is really like
- Three seminars:
 - 1. May 3, 10-12
 - 2. May 22, 13-15
 - 3. May 29, 10-12
- Instructions available from course homepage



Seminars

- **Purpose:**
 - Practice finding and reading relevant research papers
 - Practice critical thinking
 - Practice analyzing astronomical data
 - Practice scientific creativity
 - Practice communication skills
 - Practice working in a team
- **What if you cannot attend the seminars?**
 - Have to hand in written report instead (→ more work!)

Seminar I: Strange galaxy

- **Grade:** Pass/fail
- Puzzle-solving game aiming to teach you about observational techniques in extragalactic astronomy
- **Preparation:** Read section 1.3-1.4 in textbook



Seminar I: Strange Galaxy

General instructions

This document provides instructions for the first of the three seminars forming part of the examination for the course *Physics of Galaxies* in 2017. This is an exercise with game-like mechanics that aims to provide insight into some of the observational techniques commonly used in extragalactic astronomy.

In the *Strange galaxy* game, you will be assigned to different teams and presented with a sequence of observational problems to solve. Each such scenario involves a set of observational data on an unusual, poorly understood galaxy or other astronomical object. The objective is to understand the nature of object by gathering additional information through carefully chosen follow-up measurements. There are many observational methods that can be applied in each situation (optical monitoring, ultraviolet spectroscopy, high-resolution near-infrared imaging etc.) and they all provide slightly different clues to the

Seminar II: An amazing discovery / Crackpot?

- **Grade:** Pass/fail
- Role-playing exercise
- **Preparation:**
 - Study the two scenarios in the instructions
 - Read the material available in the student portal

Seminar II: An amazing discovery / Crackpot?

General instructions

This document provides instructions for the second of the three seminars forming part of the examination for the course *Physics of Galaxies* in 2017. This is a role-playing exercise that will cast you into situations that scientists (and especially astronomers) frequently encounter, yet in general tend to be rather poorly prepared for.

The point of this exercise is to:

- Practice reading research papers, press releases and other scientific texts in the field of extragalactic astronomy

Seminar III: The most distant galaxies

- **Grade:** Fail, 3, 4, 5
- **Preparation:**
 - Read suggested papers + others
 - Answer questions + analyze dataset
 - Prepare to present answers and results in class

Seminar III: The most distant galaxies

General instructions

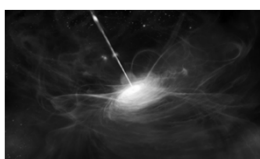
This document provides preparation instructions for third of the three seminars forming part of the examination for the course *Physics of Galaxies* in 2017. The topic of this seminar is *The most distant galaxies*.

Galaxies are being detected at ever-increasing redshifts, and as of 2017, a number of photometric galaxy candidates have been claimed at $z \approx 10-12$, i.e. at about 300-500 Myr after the Big Bang. So far, only one such object has been confirmed through spectroscopy (at $z \approx 11.1$), but the race to detect even more of these primordial galaxies is on.

The point of this exercise is to:

Database exercise ("lab")

- Introduction to exercise in lecture 6
- Complete individually and hand in report no later than June 7
- **Grade:** Fail, 3, 4, 5



Intermission: What are you looking at?



Schedule I

Complete schedule on course homepage!

- 8 Lectures:
 - March 23, 13—15
 - March 28, 13—15
 - March 30, 15—17
 - April 7, 15—17
 - April 19, 10-12
 - April 21, 13—15 ← Includes introduction to database exercise
 - April 25, 15—17
 - April 26, 10—12
- 3 Exercise sessions:
 - April 20, 13—15
 - April 27, 13—15
 - May 17, 10—12

Schedule II

- 3 seminars:
 1. May 3, 10-12
 2. May 22, 13-15
 3. May 29, 10-12
- Oral presentations of literature exercises
 - May 30, 15—17 + additional date?



Grades

- Final grade will be the mean grade from:
 - Seminar 3
 - Written report on literature exercise
 - Oral presentation of literature exercise
 - Report from database exercise
 - Hand-in exercises
- No final grade will be computed until you have reached a passing grade (3 or higher) for each of these
- Please note that you also need a passing grade from the three exercise sessions and seminar 1 & 2 to complete the course

Grades – example

- 1) Seminar 3
Grade: 4
- 2) Written report on literature exercise
Grade: 4
- 3) Oral presentation on literature exercise
Grade: 3
- 4) Report on computer exercise
Grade: 5
- 5) Hand-in exercises
Grade: 3

Calculate mean grade: $(4 + 4 + 3 + 5 + 3) / 5 = 3.8 \approx 4$
Final grade: 4:

How much time will I have to spend on this course?

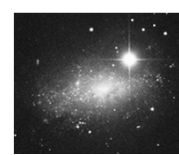
My estimates:

- Attending lectures, exercise sessions, seminars etc. ≈ 0.75 week
- Reading the textbook ≈ 1.25 weeks
- Preparing for exercise sessions ≈ 0.5 week
- Preparing for seminars (mostly seminar III) ≈ 1 week
- Computer exercise ≈ 1 week
- Literature exercise (written report + oral presentation) ≈ 1.5 weeks
- Hand-in problems ≈ 0.5 week

Sum: 6.5 weeks, i.e. ≈ 10 hp

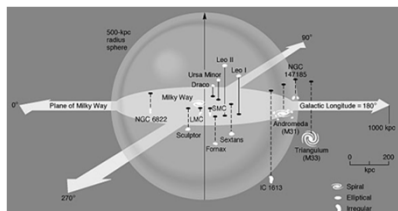
Course Outline

- Lecture 1:
 - Introduction
 - Historical Background
 - Galaxy Classification
 - The Cosmological Framework



Course Outline

- Lecture 2:
 - The Astronomical Distance Scale
 - The Milky Way
 - The Local Group



Course Outline

- Lecture 3:
 - Dark matter in galaxies
 - The dark halo
 - Subhalos
 - Mass-to-light ratios
 - Baryon fractions



Course Outline

- Lecture 4:
 - Disk galaxies
 - Elliptical galaxies



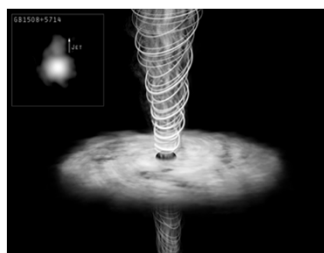
Course Outline

- Lecture 5:
 - Star formation
 - Population synthesis
 - Galaxy spectra
 - The interstellar medium
 - The cosmic star formation history



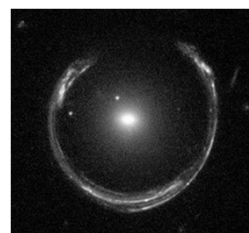
Course Outline

- Lecture 6:
 - Black holes
 - Active galaxies:
 - Quasars
 - Blazars
 - Seyfert Galaxies
 - Radio Galaxies
 - Introduction to database exercise



Course Outline

- Lecture 7:
 - Galaxy groups
 - Galaxy clusters
 - Gravitational lensing

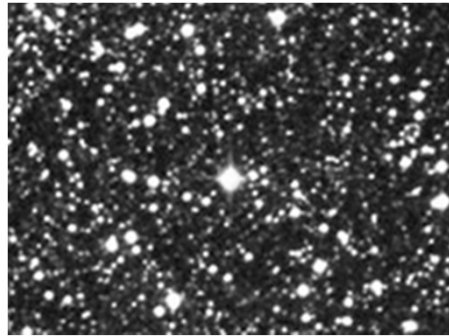


Course Outline

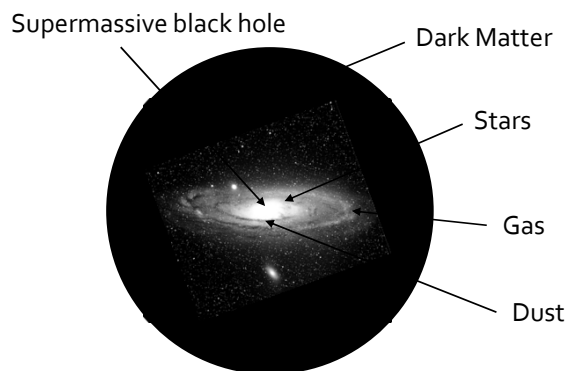
- Lecture 8:
 - The high-redshift Universe
 - Cosmic reionization
 - The first stars and galaxies



Intermission: What are you looking at?



The Anatomy of Galaxies



Historical Background: The Milky Way

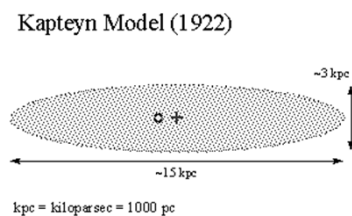


Historical Background: The Milky Way



- The "Herschel Universe" (late 1700s): Sun almost in the centre of Milky way
- Dust obscuration towards centre of the Milky Way (left side of figure) not accounted for

Historical Background: The Milky Way

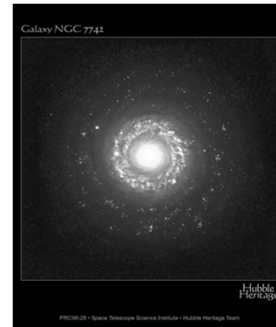


- Sun offset from centre
- Dust obscuration still not accounted for → wrong scale

Historical Background: Other Galaxies

- Mid-1800s: William Parsons (Lord Rosse) discovers spiral structure in nebulae
- 1912: Henrietta Leavitt discovers period-luminosity relation for Cepheids
- 1920s – The Great Debate
 - Shapley (local objects) VS Curtis (outside Milky Way)
 - Outcome: Spiral Nebulae are external galaxies
- 1929 – Expansion of the Universe (Hubble's law)

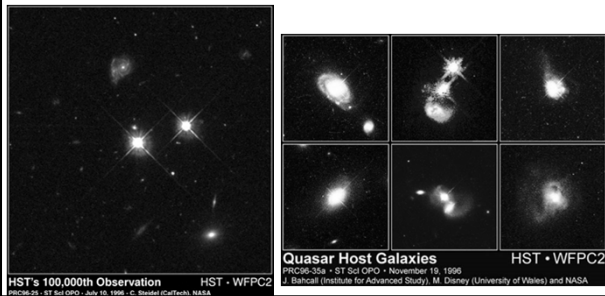
Historical Background: Active Galaxies



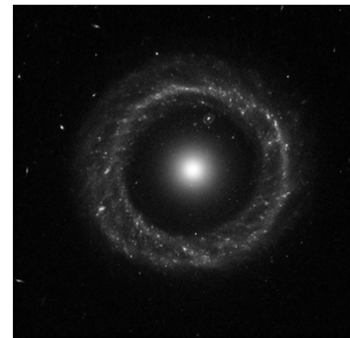
- 1943 - Seyfert Galaxies

Historical Background: Quasars/QSO

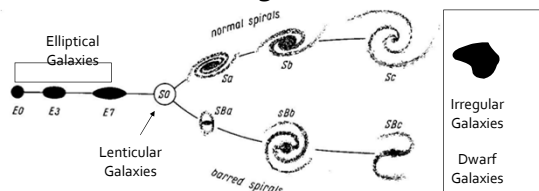
- 1960s – Radio Galaxies, Quasars
- Quasi-Stellar Object: QSO, Quasar



Intermission: What are you looking at?



Galaxy Classification The Hubble Tuning Fork



Other famous classification schemes:

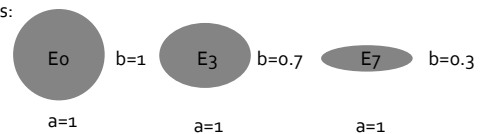
- de Vaucouleur
- van den Bergh
- Vorontsov-Velyaminov

Galaxy Classification

Elliptical galaxies

- Type: E_n , $n = 10(a - b) / a$.
- Major and minor axes: a and b
- → E0 circular, E7 galaxies the most flattened.

Examples:



Galaxy Classification

Lenticular galaxies

- Disk and central bulge, but no spiral arms
- SBo if barred, So otherwise

Examples:



So



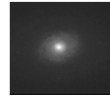
Edge-on



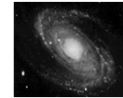
Face-on

Galaxy Classification

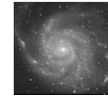
Normal Spirals



Sa

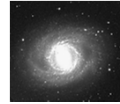


Sb

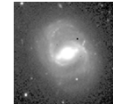


Sc

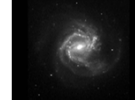
Barred Spirals



SBa



SBb

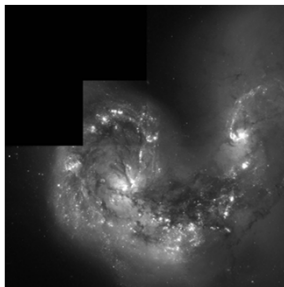


SBc

- | | |
|---------------------------------------------------------------------------------------------------------------------------------------------------|----------------------------------------------------------------------------------------------------------------------------------------------------|
| <ul style="list-style-type: none"> • Large bulges • Tightly wound spiral arms • Few star-forming regions in arms | <ul style="list-style-type: none"> • Small bulges • Loosely wound spiral arms • Many star-forming regions in arms |
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Galaxy Classification

Irregular galaxies (I)



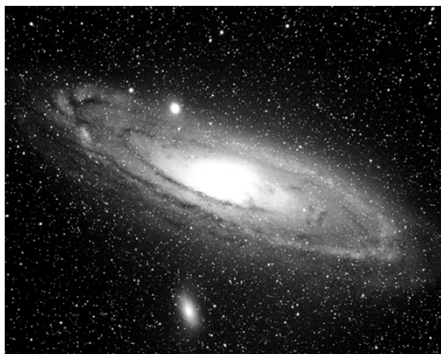
Galaxy Classification

Dwarf galaxies (dE, dSph, dI...) – Low-luminosity objects

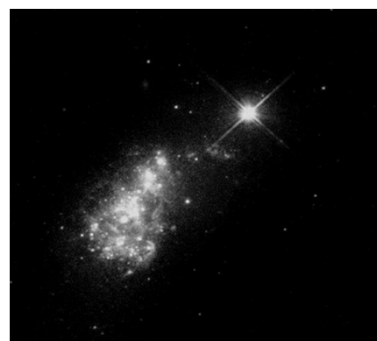


Dwarf Elliptical Galaxy M32 HST • STIS
NASA and T. Brown (NASA Goddard Space Flight Center)
STScI PR099-40

Morphological Type?



Morphological Type?

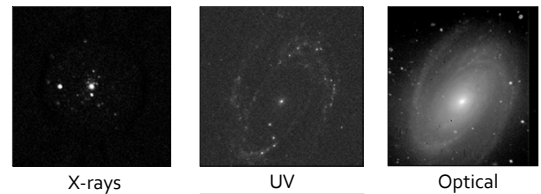


What is the Point of Morphological Classification?

Hubble class correlates with:

- Gas content
- Dust content
- Star-forming properties
- Spectrum
- Metallicity

Morphological Complications



Same galaxy (M81)
at different
wavelengths!

Morphological Complications

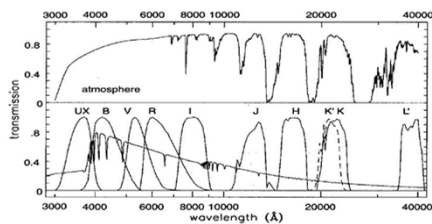
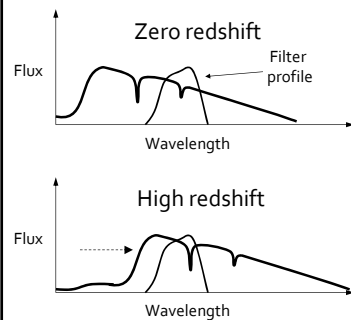


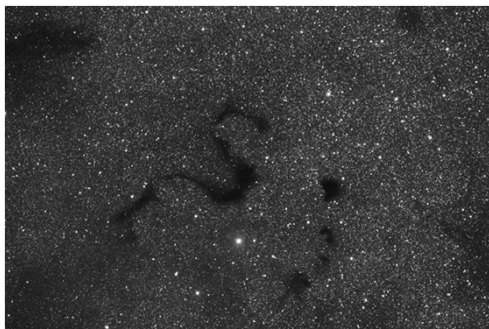
Figure 1.7 Above, atmospheric transmission in the optical and near-infrared. Below, flux F_λ of a model A0 star, with transmission curves $T(\lambda)$ for standard filters from Bessell, PASP 102, 1181; 1990. UX is a version of the U filter that takes account of atmospheric absorption. For $JHK'KL'$, $T(\lambda)$ is for transmission through the atmosphere and subsequently through the filter.

Morphological Complications

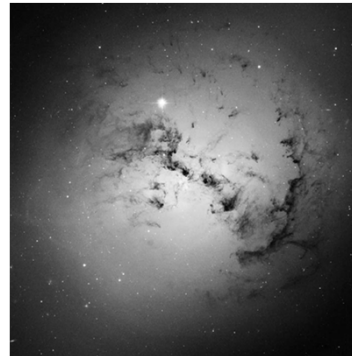


- Same filter probe different parts of spectrum at high and low redshifts
- Two galaxies with identical morphologies at a given rest wavelength may appear to have different morphologies

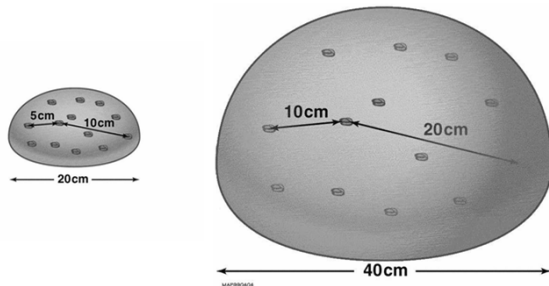
Intermission: What are you looking at?



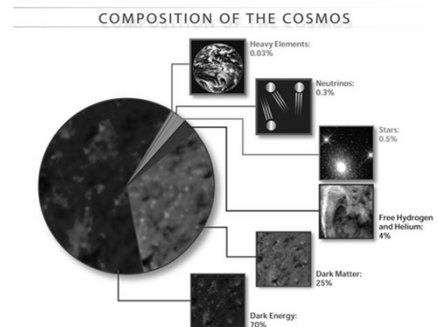
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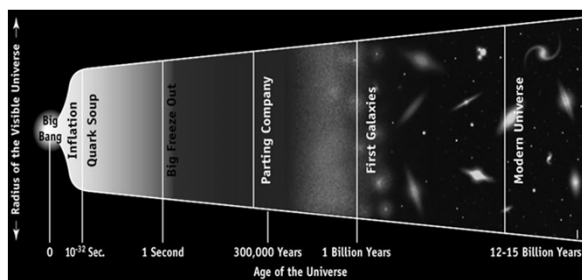
The Cosmological Framework I



The Cosmological Framework II



The Cosmological Framework III



The Cosmological Framework IV

- $\Omega_i = \rho_i / \rho_c$
- ρ_c = critical density of the Universe
- $\Omega_{\text{Tot}} \approx 1.0$
- $\Omega_{\text{Baryons}} \approx 0.04$
- $\Omega_M \approx 0.3$
- $\Omega_\Lambda \approx 0.7$