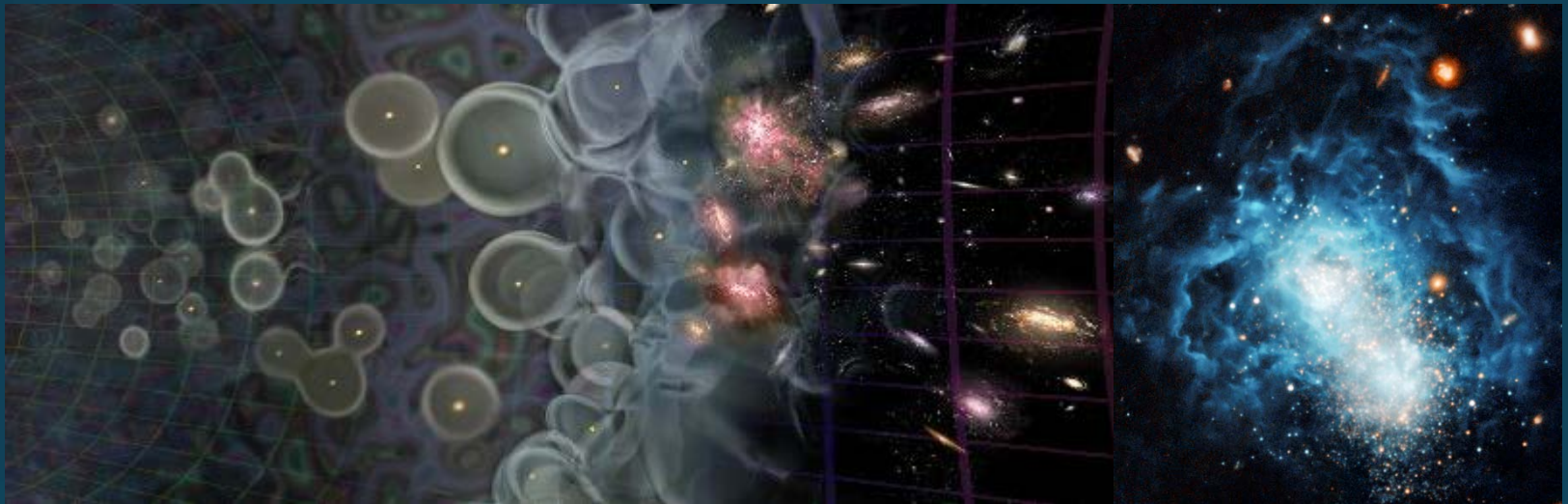


Constraining the escape of ionizing photons from $z > 6$ galaxies with JWST

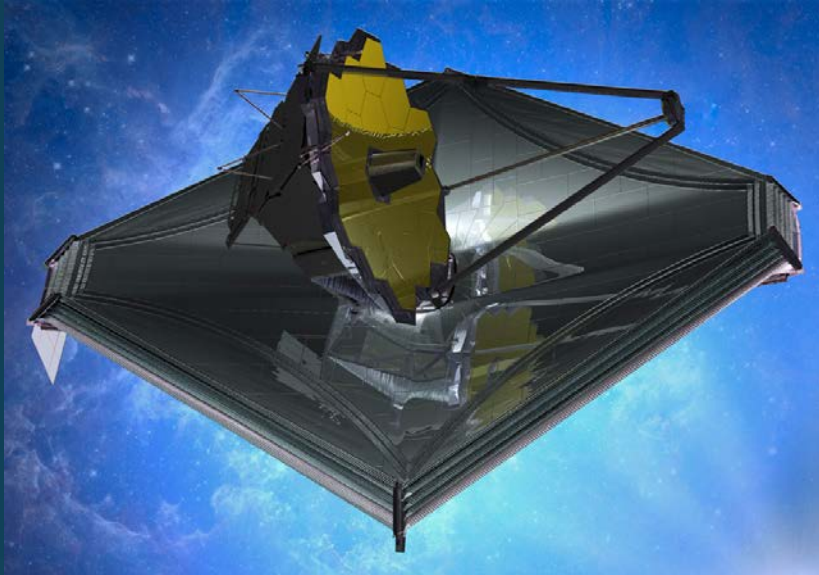


Erik Zackrisson

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Christian Binggeli, Kristian Finlator, Nick Gnedin,
Akio K. Inoue, Hannes Jensen, Genoveva Micheva,
Jan-Pieter Paardekooper, Ikkoh Shimizu

Punchline



New way to study the role of galaxies in the reionization of the Universe:

JWST/NIRSpec measurements of the escape of ionizing radiation from galaxies at redshift $z \approx 6-9$

Zackrisson et al. 2013, ApJ, 777, 39

What caused cosmic reionization?

Studies of the intergalactic medium (e.g. $\text{Ly}\alpha$, CMBR) can reveal *when* reionization happened, but where did the ionizing photons come from?

Star-forming galaxies at $z > 6$ are the prime suspects, but how do we prove that they did it?



CMBR

Reionization
($z \approx 6-14$)

Today

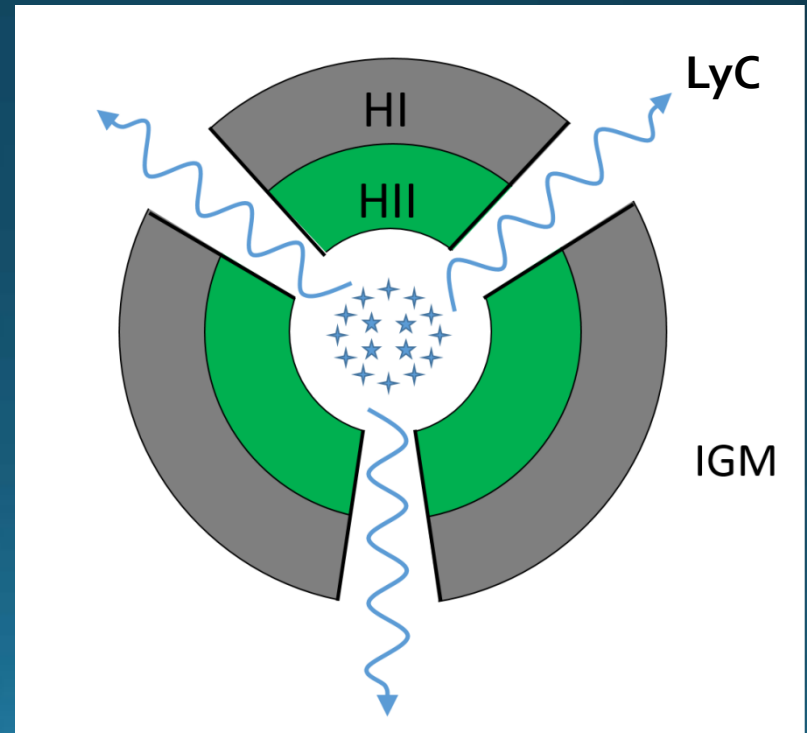
S.G. Djorgovski et al. & Digital Media Center, Caltech

Requirements for galaxy-dominated reionization

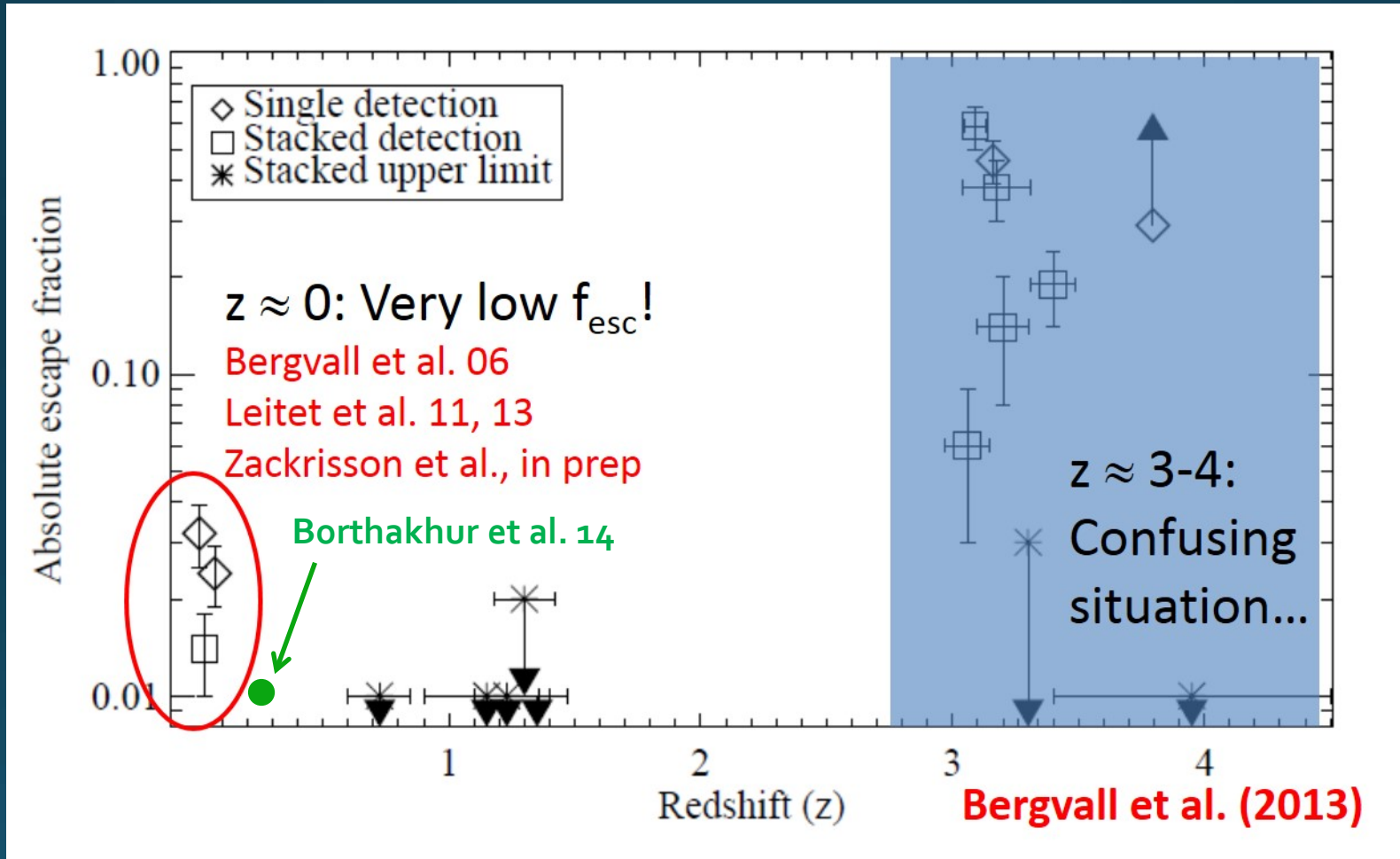
1. Lots of star-forming galaxies at redshifts $z > 6$



2. Escape of ionizing photons (Lyman continuum) from galaxies into the IGM

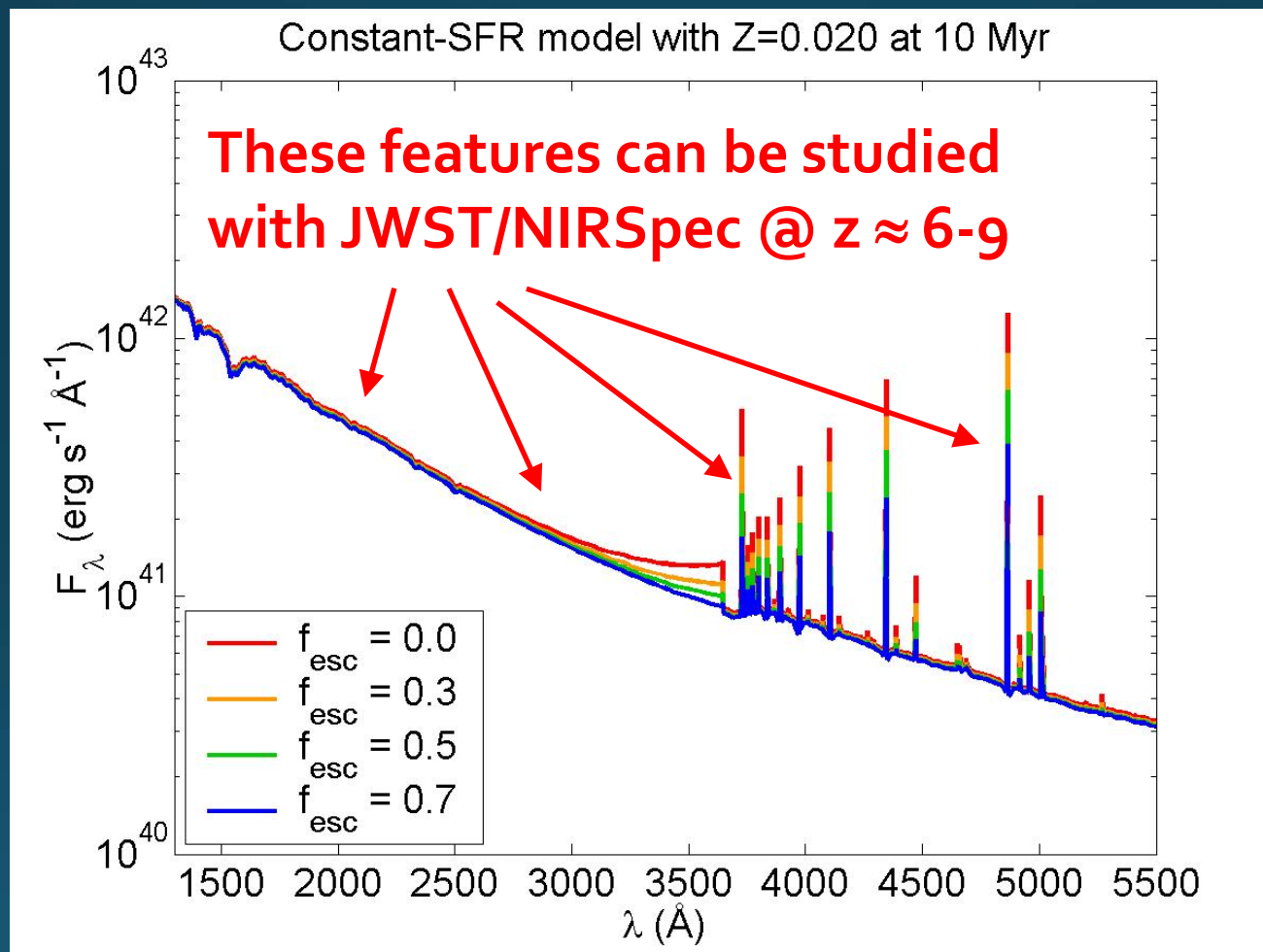


LyC leakage – direct detections



Serious problem: At $z > 4-5$, the IGM opacity prevents direct detection of escaping Lyman continuum (Inoue & Iwata 2008, Inoue+14)

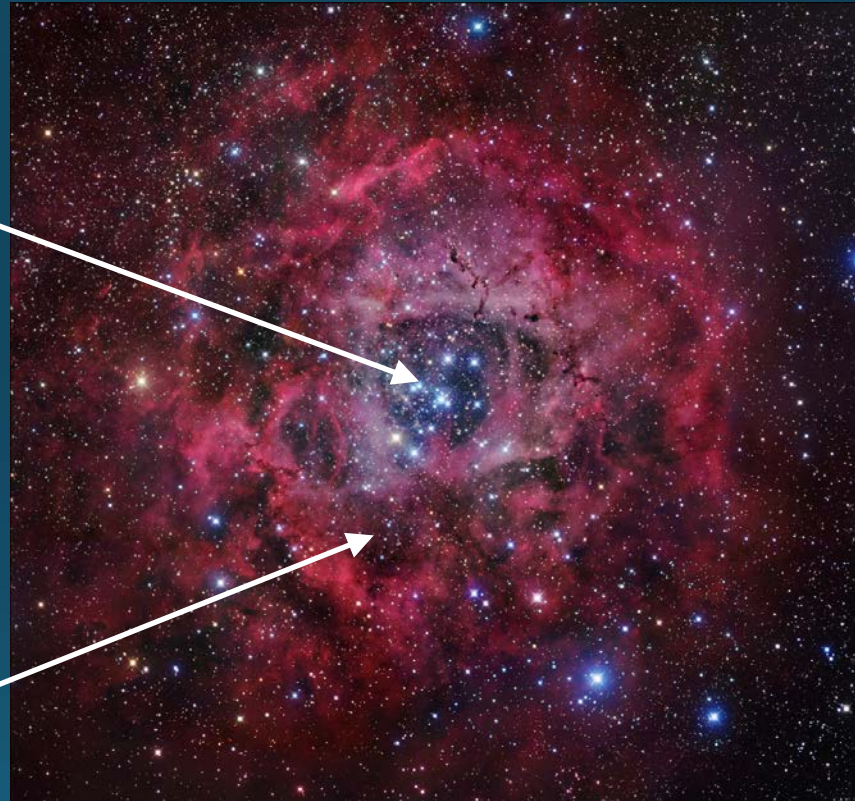
Indirect signatures of LyC leakage



How does this work?

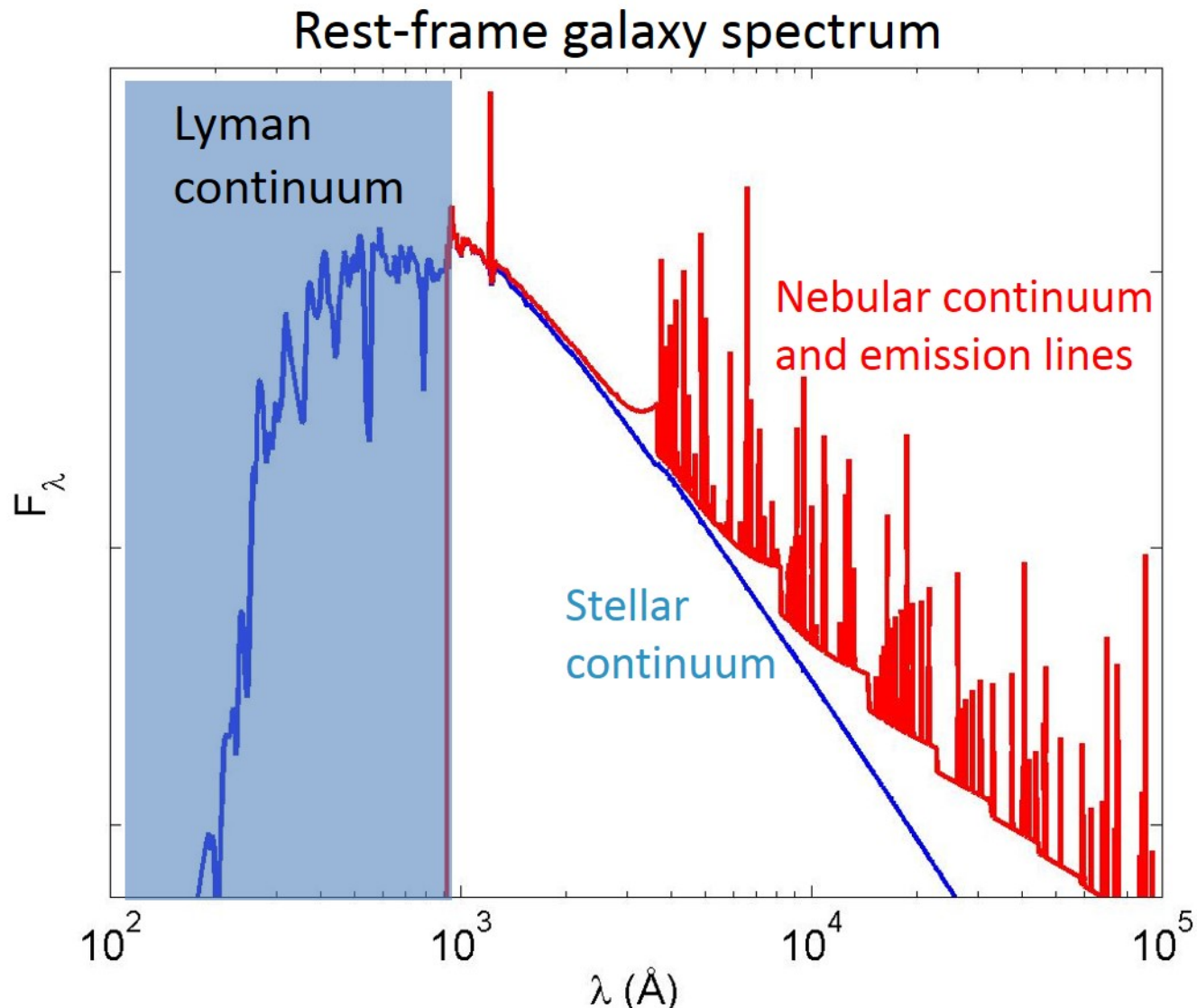
Young stars

Lyman continuum
photons from young
stars captured by gas
→ **Nebular emission**



If some of the LyC photons escape without
ionizing the ISM → **Less nebular emission!**

Stellar and nebular SEDs





- Pop I, II, III stars + EMP stars
- Nebular emission (Cloudy) + dust
- Rest-frame SEDs (far-UV to near-IR)
- HST/Spitzer/JWST fluxes @ $z=0-15$

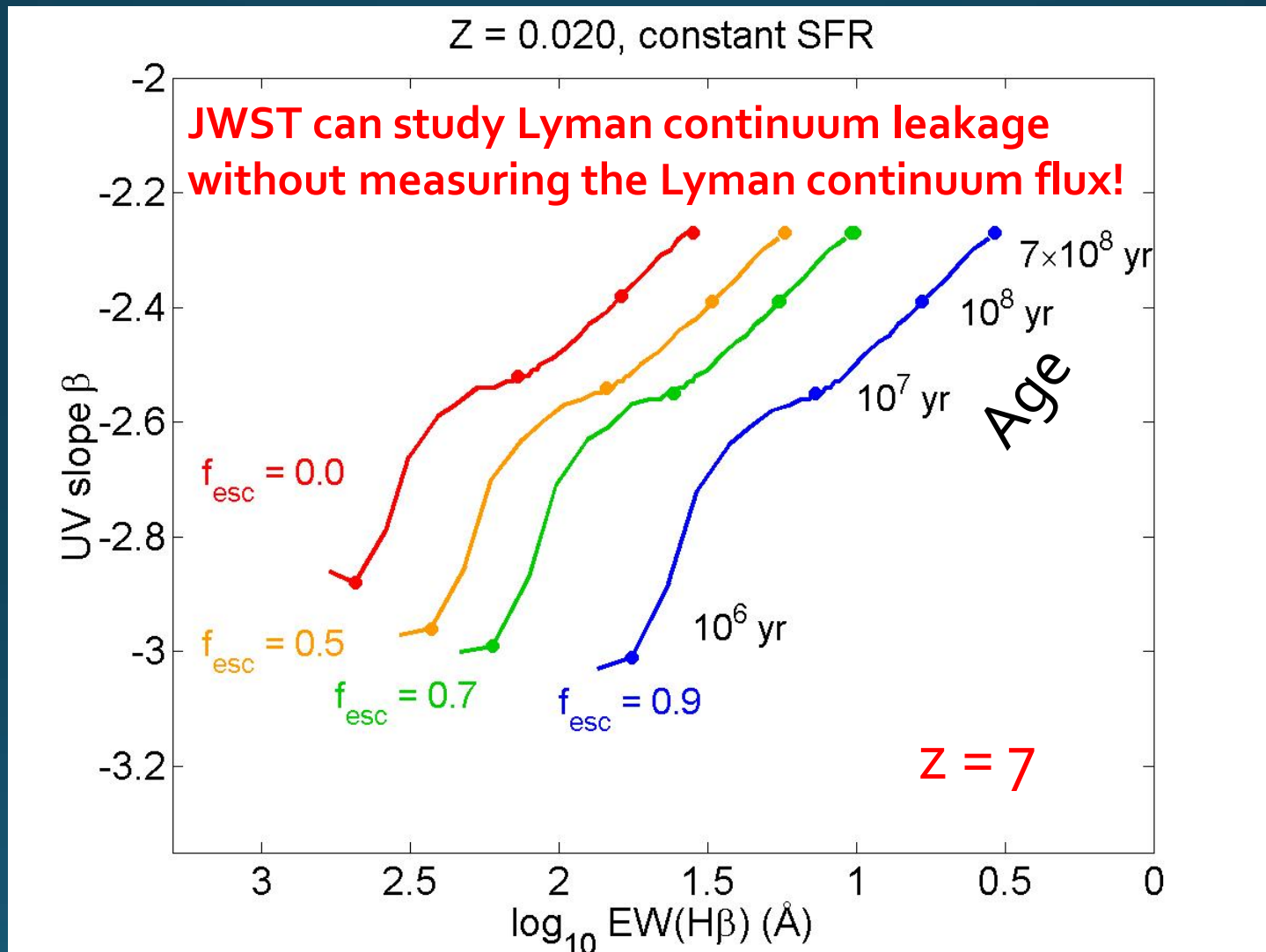
The **ggdrasil** code

A spectral synthesis model for the first galaxies

Model grids available at: www.astro.uu.se/~ez

Zackrisson et al. 2011, ApJ, 740, 13

Simple diagnostics: UV slope & H β



How well can we measure f_{esc} ?



LYman Continuum ANalysis project:

Galaxy simulations + Yggdrasil + observational errors →

Mock spectra of high- z galaxies with Lyman continuum leakage

Simulations suites:

CROC (Gnedin 2014)

FiBY (Paardekooper et al. 2015)

Finlator et al. (2013)

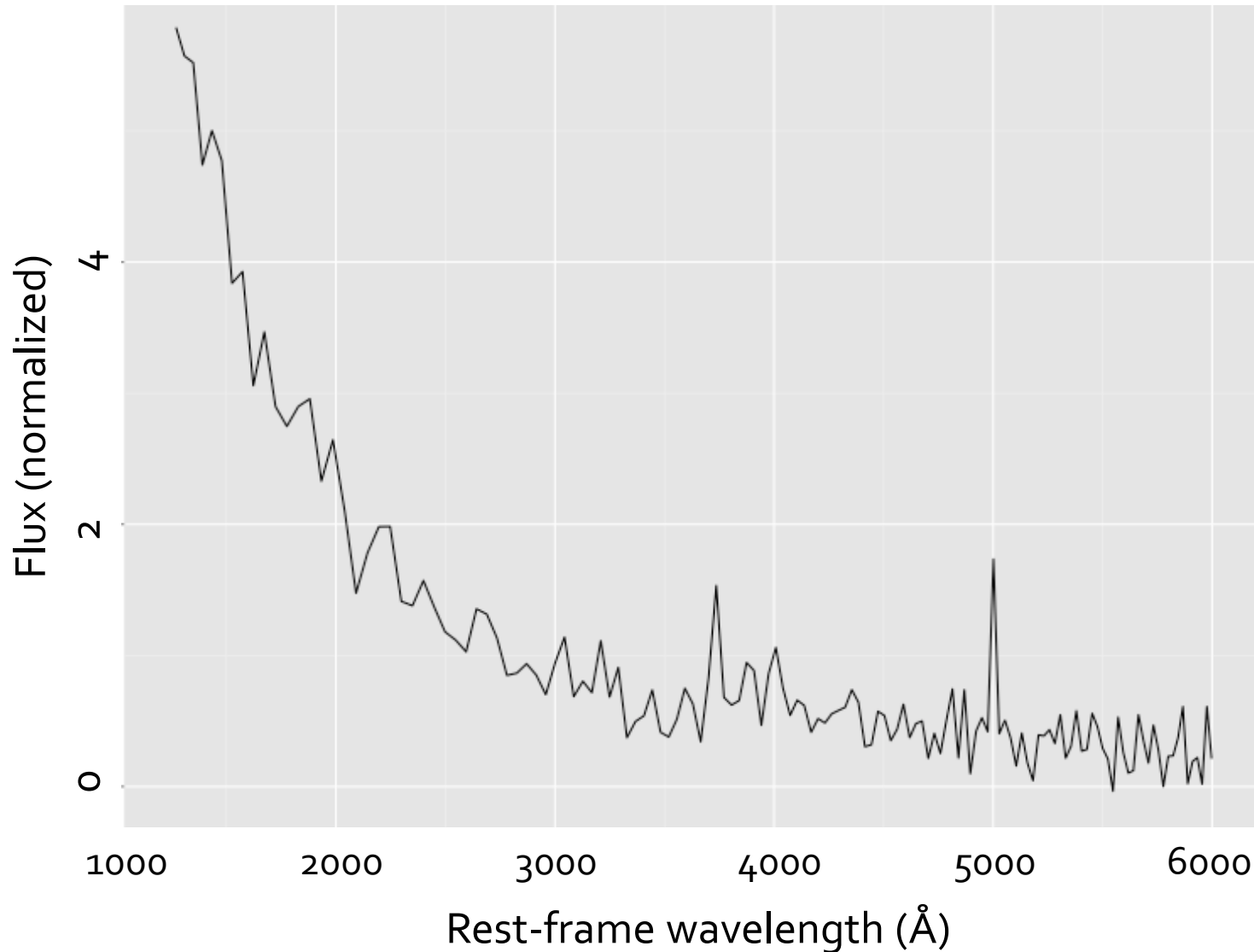
Shimizu et al. (2014)

Highly realistic SEDs for high-redshift galaxies publicly available:

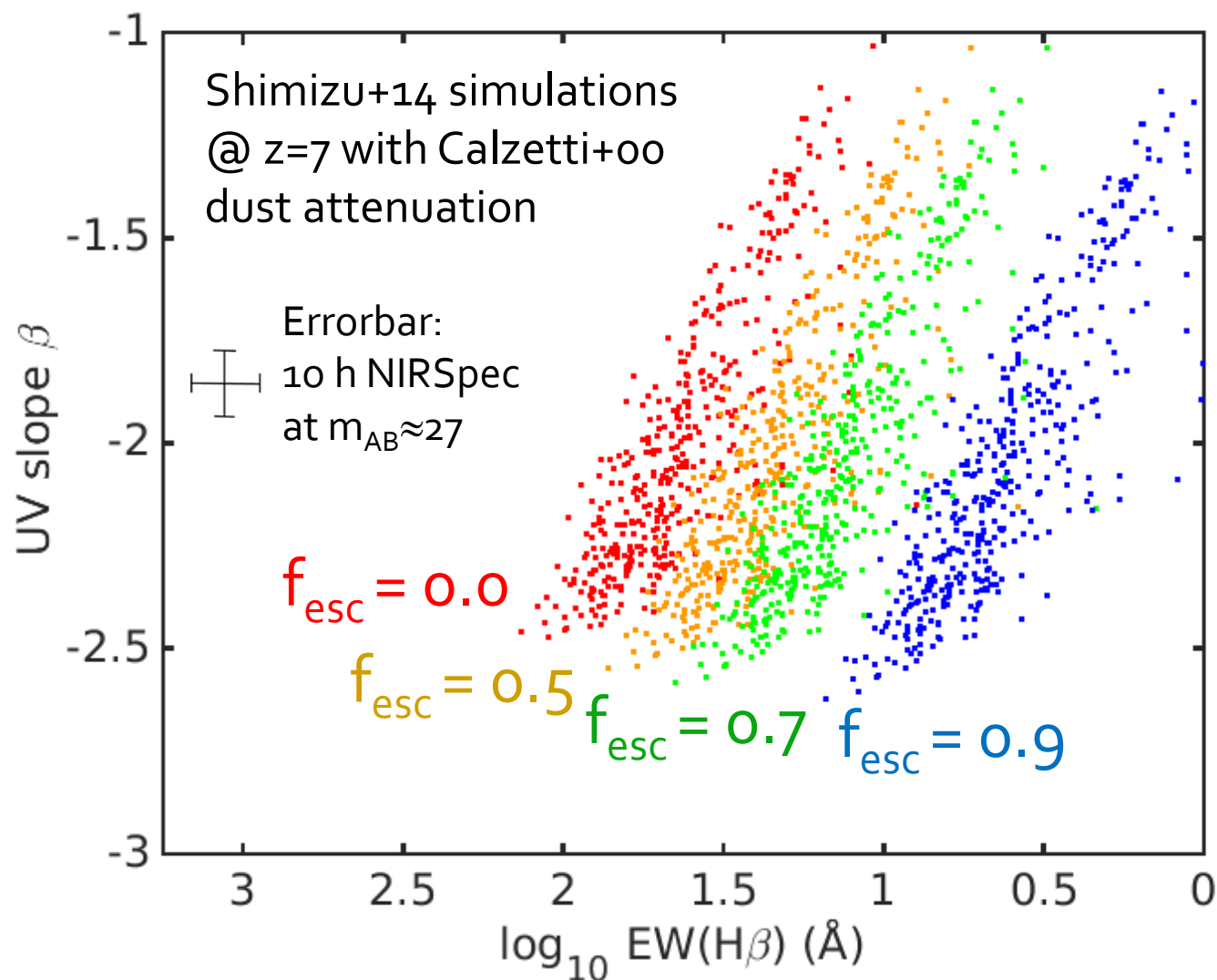
www.astro.uu.se/~ez/lycan/lycan.html

Example of mock SED from LYCAN

3h NIRS pec, R=100 spectrum of $m_{AB} \approx 27$ galaxy at $z=7$

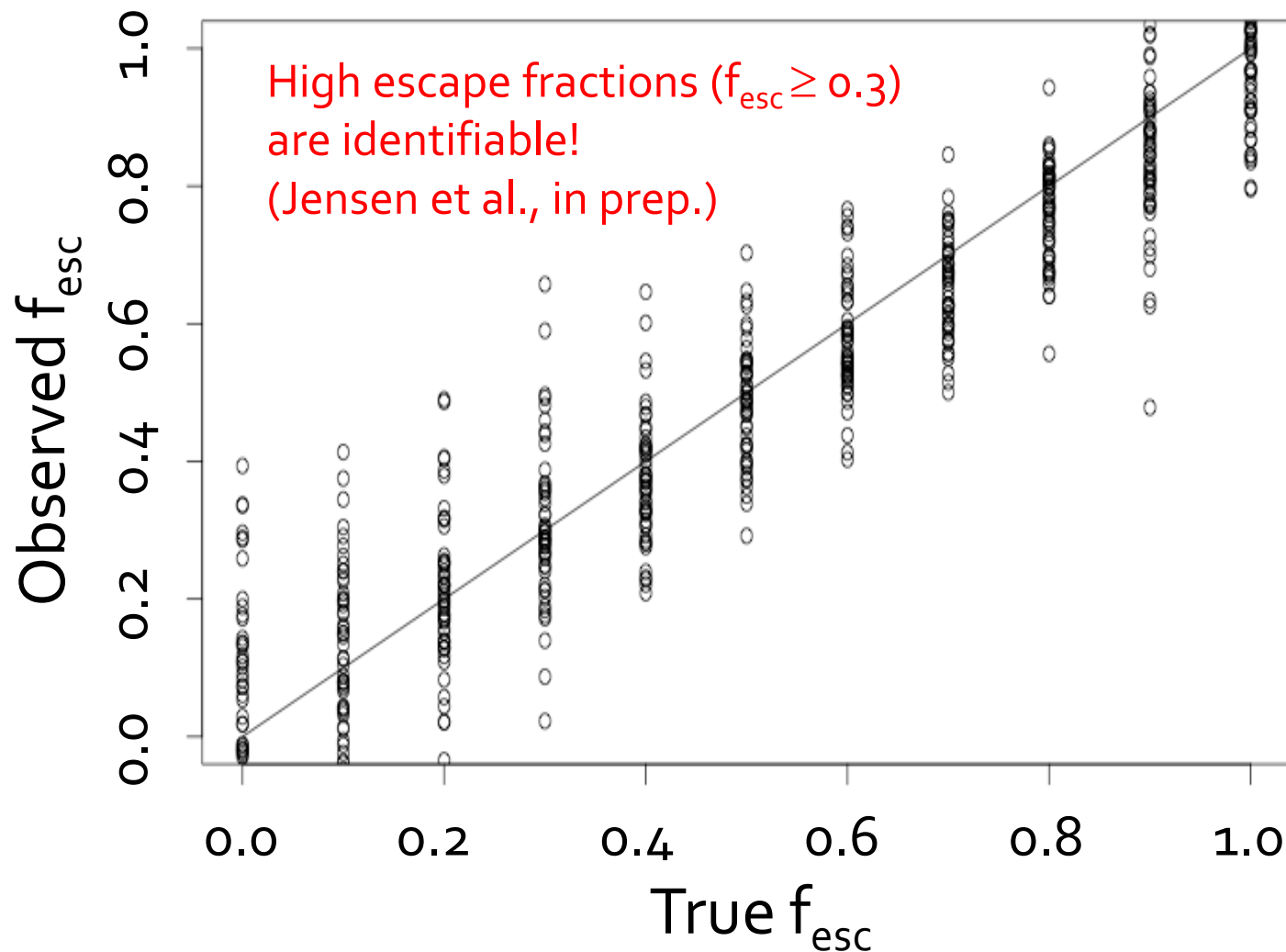


Diagnostics based on realistic SEDs



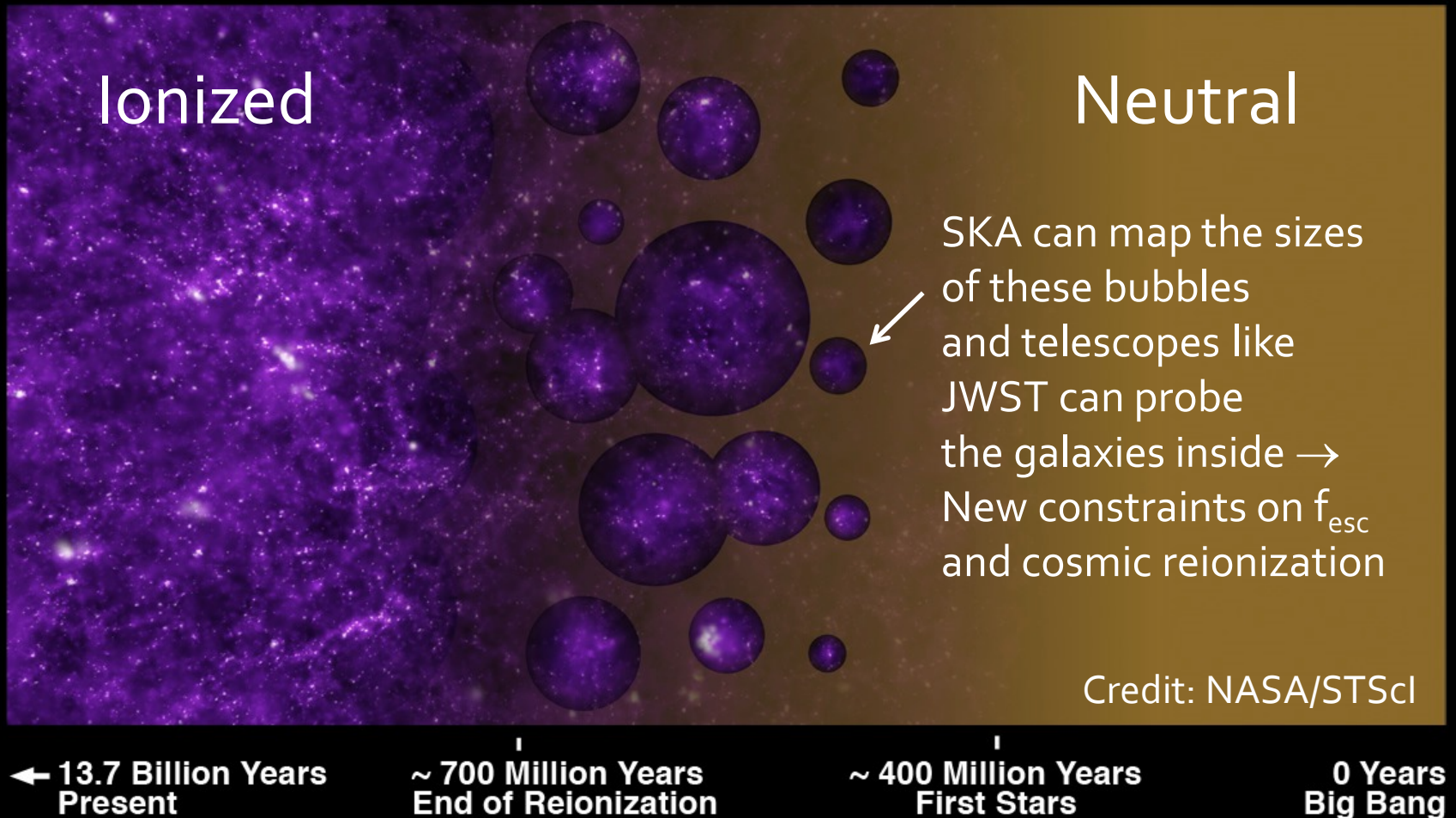
SED analysis using machine learning algorithms

3h NIRS pec, R=100 spectra of $m_{AB} \approx 26-27$ galaxies at $z=7$



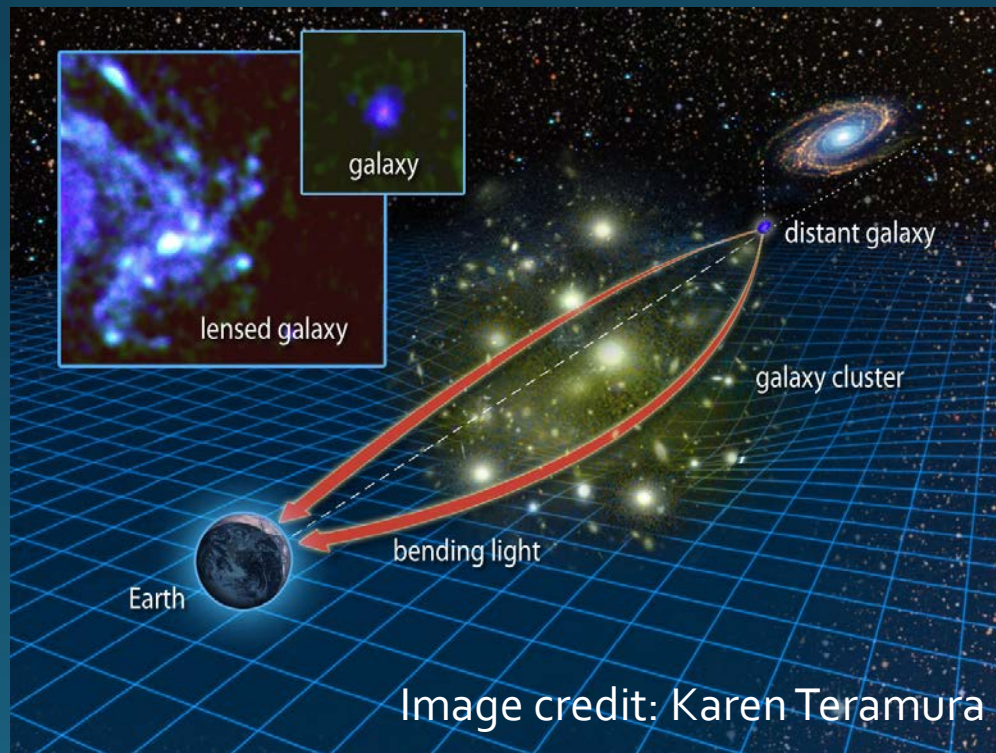
Ongoing work: Synergies with the Square Kilometer Array

Time Since The Big Bang



Lensed galaxies make the best targets

In <10 h, NIRSpec can get a sufficiently good spectrum for a $z \approx 7$ galaxy with stellar mass $\sim 10^7$ Msolar ($M_{1500} \approx -16$) if the gravitational magnification is $\mu \approx 30$



Summary

- The escape of ionizing photons from galaxies at $z > 6$ is crucial for galaxy-dominated reionization
- JWST/NIRSpec can constrain the escape of ionizing radiation from galaxies up to $z \approx 9$
- Project can go piggy-back on any NIRSpec survey of $z > 6$ galaxies, but lensed galaxies make the best targets
- Publicly available model SEDs for high- z galaxies:
www.astro.uu.se/~ez/lycan/lycan.html

